

# Universality in 2D CFT and 3D Gravity

## How to Toast your Bagels Perfectly Every Time

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# Outline

- 1 Modular Invariance in 2D CFT**
- 2 Universal Features of the Spectrum
  - The Thermal Partition Function
  - Thermodynamics and Spectrum
- 3 3D Gravity: Biting Into the Bagel
  - The Classical Bulk Solutions
  - The Hawking–Page Transition
- 4 3D Gravity from 2D CFT

# The Partition Function in QFT

Consider an arbitrary QFT on  $S_L^1 \times \{\text{time}\}$  at temperature  $T = \frac{1}{\beta}$  with Hamiltonian  $H$ . The Hilbert space  $\mathcal{H}$  consists of states on  $S_L^1$ , so  $H$  has a discrete spectrum  $\{E_n\}$  bounded below. The partition function is

$$Z_L(\beta) \equiv \text{Tr}_{\mathcal{H}}(e^{-\beta H}) = \sum_n \langle n | e^{-\beta H} | n \rangle = \sum_n e^{-\beta E_n}. \quad (1.1)$$

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We can always write  $Z_L(\beta)$  as a Euclidean path integral on the bagel:

$$\begin{aligned} Z_L(\beta) &= \int_{T^2(L,\beta)} \mathcal{D}\phi e^{-S[\phi]} = \beta \left[ \begin{array}{c} \text{---} \\ \square \\ \text{---} \\ L \end{array} \right] = \beta \left[ \begin{array}{c} \square \\ \text{---} \\ L \end{array} \right] = \\ &= \int_{T^2(\beta,L)} \mathcal{D}\phi e^{-S[\phi]} = Z_\beta(L). \end{aligned} \quad (1.2)$$

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This relation is completely general. It relates two *different* theories with the same Euclidean action but on spatial manifolds of different sizes.

# The Partition Function in CFT

If our theory is scale-invariant, then  $Z_L(\beta) = L_{\alpha L}(\alpha\beta)$ , and only the “bagel aspect ratio”  $\frac{\beta}{L}$  matters. Thus we set  $L = 2\pi$ . It follows that

$$Z(\beta) \equiv Z_{2\pi}(\beta) = Z_{\beta}(2\pi) = Z_{2\pi}\left(2\pi \cdot \frac{2\pi}{\beta}\right) = Z\left(\frac{4\pi^2}{\beta}\right). \quad (1.3)$$

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A 2D CFT is **modular invariant** if one has

$$Z(\beta) = Z\left(\frac{4\pi^2}{\beta}\right).$$

**The UV spectrum and thermodynamics are determined by the IR.**  
But the IR is determined largely by the ground state, which is *universal*.

More generally,  $\beta$  is replaced by  $\tau = \frac{i\beta}{2\pi} \in \mathbb{C}$ , and in fact  $Z(\tau) = Z(-\frac{1}{\tau})$ .  
Even more generally,  $Z(\tau)$  is invariant under  $\text{PSL}(2, \mathbb{Z})$  transformations.

# Operators and their Dimensions

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- The central charge is  $c = c_L + c_R$ , and we assume that  $c_L = c_R$ .
- The dimension of  $\mathcal{O}$  is  $\Delta = h + \bar{h}$ , while its spin is  $J = h - \bar{h}$ . We will treat only the case  $J = 0$  of zero angular potential.
- In radial quantization, the Hamiltonian is a dilation operator,  $H = D - \frac{c}{12}$ , shifted by the Casimir energy  $E_0 = -\frac{c}{12} \frac{2\pi}{L}$ .
- The eigenvalues of  $H$  are  $E$ ; the eigenvalues of  $D$  are  $\Delta$ .
- The ground state  $|0\rangle$  corresponds to the identity operator  $\mathbb{1}$ , which is the unique operator with  $\Delta_0 = 0 \iff E_0 = -\frac{c}{12}$ .

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Including degeneracies,  $Z(\beta) = \sum_E \rho(E) e^{-\beta E} = \sum_{\Delta} \rho(\Delta) e^{-\beta(\Delta - \frac{c}{12})}$ ,

where the density of states  $\rho(E) = e^{S(E)}$  is related to the entropy.

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# The Thermal Partition Function

**Goal:** to study the thermodynamics of 2D CFTs at high temperature.

High temperature:  $\beta \rightarrow 0$ . Then  $Z(\beta) \rightarrow \sum_E \rho(E) e^{-0E} \sim N \sim e^S$ .

Low temperature:  $\beta \rightarrow \infty$ . Then  $Z(\beta) \rightarrow \sum_E e^{-\infty E} = e^{-\beta E_0} = e^{\frac{\beta c}{12}}$ .

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Low temperature:  $\beta \rightarrow \infty$ . Then  $Z(\beta) \rightarrow \sum_E e^{-\infty E} = e^{-\beta E_0} = e^{\frac{\beta c}{12}}$ .

These two regimes are linked by modular invariance. At high temperature,

$$Z(\beta) = Z\left(\frac{4\pi^2}{\beta}\right) = e^{\frac{c}{12} \cdot \frac{4\pi^2}{\beta}} = e^{\frac{\pi^2 c}{3\beta}} \implies \boxed{\log Z(\beta) = \frac{\pi^2 c}{3\beta}}. \quad (2.1)$$

This statement is *asymptotic*: it's a good enough approximation at high enough temperatures, but is silent on *when* it's valid and *how good* it is.

# Thermodynamics and Spectrum

Let's make like undergraduates and compute! The free energy  $F(\beta)$  is

$$\log Z(\beta) = \frac{\pi^2 c}{3\beta} = \beta F(\beta) \implies F(\beta) = \frac{\pi^2 c}{3\beta^2} = \frac{c}{12} \left( \frac{2\pi}{\beta} \right)^2. \quad (2.2)$$

The thermodynamic entropy  $S(\beta)$  at high temperature is obtained from

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We change to the microcanonical ensemble via  $\langle E \rangle_\beta = -\partial_\beta \log Z(\beta)$ :

$$\boxed{S(E) = 2\pi \sqrt{\frac{c}{3} E} \iff \rho(\Delta) = \exp \left[ 2\pi \sqrt{\frac{c}{3} \left( \Delta - \frac{c}{12} \right)} \right].} \quad (2.4)$$

This is the **Cardy formula**. It says that at high energies, **all unitary, modular invariant 2D CFTs have universal thermodynamics**.

# Comments on the Cardy Formula

- 1 The asymptotic behavior modular-invariant 2D CFTs is completely fixed by the vacuum state, i.e. by the mighty identity operator.



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- 2 Modular invariance means that *both* IR and UV are universal:

$$\log Z(\beta) = \frac{c}{12} \begin{cases} \frac{4\pi^2}{\beta}, & \beta \ll 2\pi, \\ \beta, & \beta \gg 2\pi. \end{cases} \quad (2.5)$$

These results are only valid away from the self-dual point  $\beta_* = 2\pi$ , but they might remind you of (!) the Hawking–Page transition.

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- 3 We've just solved the **bootstrap equation**

$$Z(\beta) = \sum_{\Delta} \rho(\Delta) e^{-\beta(\Delta - \frac{c}{12})} = e^{\frac{\pi^2 c}{3\beta}} \quad (2.6)$$

for  $\rho(\Delta)$ . The singular term on the RHS can teach us about asymptotics on the LHS: that's how the bootstrap works.

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# The Classical Bulk Solutions

**Semiclassical AdS<sub>3</sub> gravity**  $\leftrightarrow$  **CFT<sub>2</sub> with**  $c = \frac{3}{2G} \gg 1$ . The spectrum is given by  $\rho(E) \sim e^{S(E)}$ , so we focus on  $S(E)$  at leading order in  $c$ .

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Classical solutions to pure 3D gravity must be locally isometric to AdS<sub>3</sub>:

- 1 **Global AdS<sub>3</sub>** corresponds to the vacuum (i.e.  $\mathbb{1}$ ):  $E = -\frac{1}{8G} = -\frac{c}{12}$ .
- 2 **BTZ black holes** correspond to excited states (i.e. heavy primaries) and are formed from AdS<sub>3</sub> by discrete identifications:  $E = M \geq 0$ .

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These have entropy in agreement with the Cardy formula (?!?!):

$$S(E) = \frac{A}{4G} = 2\pi\sqrt{\frac{c}{3}E} \quad (A \sim \sqrt{M}) \quad (3.1)$$

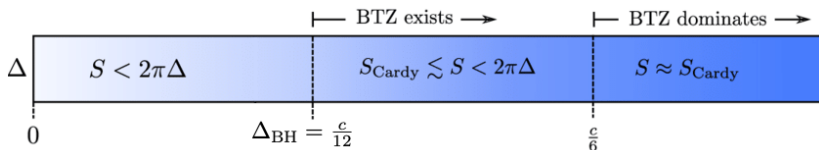
Unlike the Cardy formula, this applies for *all*  $E$ , as long as  $c \gg 1$ .

# The Classical Bulk Solutions

- 3 More generally, there are  $SL(2, \mathbb{Z})$  black holes that replace  $\beta$  by the complex modulus  $\tau \in \mathcal{F} = \mathbb{H}^2 / PSL(2, \mathbb{Z}) \subset \mathbb{C}$ .
- 4 One can apply large diffeomorphisms to excite “boundary gravitons” and create moving BHs. (In CFT, these correspond to descendants.)
- 5 Turning on matter, one finds conical defects with  $-\frac{c}{12} < E < 0$ , subleading horizonless geometries, and exotic black objects.

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# A Sparse Light Spectrum

So **pure AdS<sub>3</sub> ( $E = -\frac{c}{12}$ ) and the BTZ geometries ( $E \geq 0$ ) are the only smooth solutions of pure 3D gravity**: there's nothing in the gap.

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**Sparseness:** Every holographic 2D CFT must have a “small” number of (primary) operators in the gap  $-\frac{c}{12} < E < 0 \iff 0 < \Delta < \frac{c}{12}$ .

In fact, we require only that  $\rho(\Delta) \leq e^{2\pi\Delta}$  for  $0 \leq \Delta \leq \frac{c}{12}$ .

**E.g.** Consider  $N = c \gg 1$  free bosons. We have  $\rho(\Delta) = \exp(2\pi\sqrt{\frac{c}{3}\Delta})$  for all  $\Delta \geq 0$ . There's no gap, so this theory cannot be holographic.

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The **Hellerman bound** states that every unitary, modular invariant 2D CFT must contain a primary with dimension  $0 < \Delta_1 \leq \frac{c}{6} + 0.473695$ . The bound has since been improved, and is expected to be  $\Delta_1 < \frac{c}{12}$ .

# The Hawking–Page Transition

The spectrum in hand, we compute the partition function of 3D gravity:

$$Z_{\text{AdS}}(\beta) = \sum_E \rho(E) e^{-\beta E} = e^{\frac{\beta c}{12}} + \int_0^\infty dE \rho(E) \exp\left(2\pi\sqrt{\frac{c}{3}E} - \beta E\right).$$

This integral has a saddle point at  $E_* = \frac{\pi^2 c}{3\beta^2}$ , so to leading order in  $c$ ,

$$Z_{\text{AdS}}(\beta) = e^{\frac{\beta c}{12}} + e^{\frac{\pi^2 c}{3\beta}} \approx \max\left\{e^{\frac{\beta c}{12}}, e^{\frac{\pi^2 c}{3\beta}}\right\}. \quad (3.2)$$

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The thermodynamics is then exactly the same as before:

$$\beta F(\beta) = \log Z(\beta) = \frac{c}{12} \begin{cases} \frac{4\pi^2}{\beta}, & \beta < 2\pi \text{ (BTZ)}, \\ \beta, & \beta > 2\pi \text{ (tAdS)}. \end{cases} \quad (3.3)$$

The **Hawking–Page transition** occurs at  $\beta_* = 2\pi$ . At low  $T$ , the vacuum (“thermal AdS”) dominates; at high  $T$ , BTZ is dominant. Of course, “small” BTZs still exist for  $\beta > 2\pi$ , where  $E_* < \frac{c}{12}$ .

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# Gravity vs. CFT

Here's we know about  $\text{AdS}_3/\text{CFT}_2$  so far:

- All 2D CFTs match 3D gravity when  $\beta \rightarrow 0$  via Cardy's formula.
- But in the limit  $c \rightarrow \infty$ , Cardy's formula works for *all*  $\beta < 2\pi$ .

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## Theorem ( $\text{AdS}_3 = \text{CFT}_2$ )

We have  $Z_{\text{AdS}} = Z_{\text{CFT}}$  iff 3 conditions are satisfied:

- 1 The CFT is unitary and modular invariant;
- 2  $c \gg 1$ , i.e. the limit  $c \rightarrow \infty$  is taken; and
- 3 The light spectrum is sparse, meaning  $\rho(\Delta) \leq e^{2\pi\Delta}$  for  $\Delta < \frac{c}{12}$ .



# Proof of AdS/CFT

## *Proof*

On the gravity side, we classified all bulk solutions as quotients of  $\text{AdS}_3$ . We determined their energies (there's a gap!), computed the entropy and density of states, and evaluated the partition function at large  $c$ :

$$Z_{\text{AdS}}(\beta) = e^{\frac{\beta c}{12}} + e^{\frac{\pi^2 c}{3\beta}}. \quad (4.1)$$

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In the CFT, the key idea is to split  $Z(\beta)$  into a sum of its contributions from light and heavy states, and then to use **modular invariance**:

$$Z_{\text{CFT}}(\beta) = Z_{\text{L}}(\beta) + Z_{\text{H}}(\beta) = Z_{\text{L}}\left(\frac{4\pi^2}{\beta}\right) + Z_{\text{H}}\left(\frac{4\pi^2}{\beta}\right). \quad (4.2)$$

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Now at **large**  $c$ , the theory's degrees of freedom decouple. That is,

$$Z_{\text{L}}(\beta) = Z_{\text{H}}\left(\frac{4\pi^2}{\beta}\right), \quad Z_{\text{H}}(\beta) = Z_{\text{L}}\left(\frac{4\pi^2}{\beta}\right). \quad (4.3)$$

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Finally, we use **sparseness**, which guarantees that

$$e^{\frac{\beta c}{12}} \leq Z_{\text{L}}(\beta) = \sum_{\Delta} \rho(\Delta) e^{-\beta(\Delta - \frac{c}{12})} \leq e^{\frac{\beta c}{12}} \sum_{\Delta} e^{-(\beta - 2\pi)\Delta}. \quad (4.4)$$

If  $\beta > 2\pi$ , then the last factor exponentially suppresses all terms except for  $\Delta = 0$ . Thus in fact  $Z_{\text{L}}$  is dominated by the vacuum, and we have

$$Z_{\text{L}}(\beta) = e^{\frac{\beta c}{12}} \implies Z_{\text{CFT}}(\beta) = e^{\frac{\beta c}{12}} + e^{\frac{\pi^2 c}{3\beta}} = Z_{\text{AdS}}(\beta). \quad (4.5)$$

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**So** CFT =  $\sum_{\text{states}}$  (fixed channel), **while** gravity =  $\sum_{\text{channels}}$  (vacuum).

# Summary and Conclusions

- Modular invariance in 2D CFT is expressed by  $Z(\beta) = Z(4\pi^2/\beta)$ . The IR, where the vacuum lives, strongly constrains the UV.
- At high temperature,  $Z(\beta)$  is dominated by the vacuum, so the thermodynamics is universal:  $S(E) = 2\pi\sqrt{\frac{c}{3}E}$  (Cardy).

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- At high temperature,  $Z(\beta)$  is dominated by the vacuum, so the thermodynamics is universal:  $S(E) = 2\pi\sqrt{\frac{c}{3}E}$  (Cardy).
- The classical solutions of pure 3D gravity are AdS<sub>3</sub> ( $E = -\frac{c}{12}$ ) and the BTZ black holes ( $E \geq 0$ ). In particular, the spectrum is gapped.
- At large  $c$ , the partition function is  $Z(\beta) = e^{\frac{\beta c}{12}} + e^{\frac{\pi^2 c}{3\beta}}$ , and there is a Hawking–Page transition between tAdS and BTZ at  $\beta_* = 2\pi$ .



# Summary and Conclusions

- Modular invariance in 2D CFT is expressed by  $Z(\beta) = Z(4\pi^2/\beta)$ . The IR, where the vacuum lives, strongly constrains the UV.
- At high temperature,  $Z(\beta)$  is dominated by the vacuum, so the thermodynamics is universal:  $S(E) = 2\pi\sqrt{\frac{c}{3}E}$  (Cardy).
- The classical solutions of pure 3D gravity are AdS<sub>3</sub> ( $E = -\frac{c}{12}$ ) and the BTZ black holes ( $E \geq 0$ ). In particular, the spectrum is gapped.
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- Any holographic CFT must be unitary and modular invariant, have large central charge, and satisfy  $\rho(\Delta) \leq e^{2\pi\Delta}$  for  $\Delta \in [0, \frac{c}{12}]$ .
- Under these conditions, the validity of the Cardy formula extends to all  $\beta$ , and moreover we have  $Z_{\text{AdS}}(\beta) = Z_{\text{CFT}}(\beta)$ .

**Thank you for listening! Any questions?**